

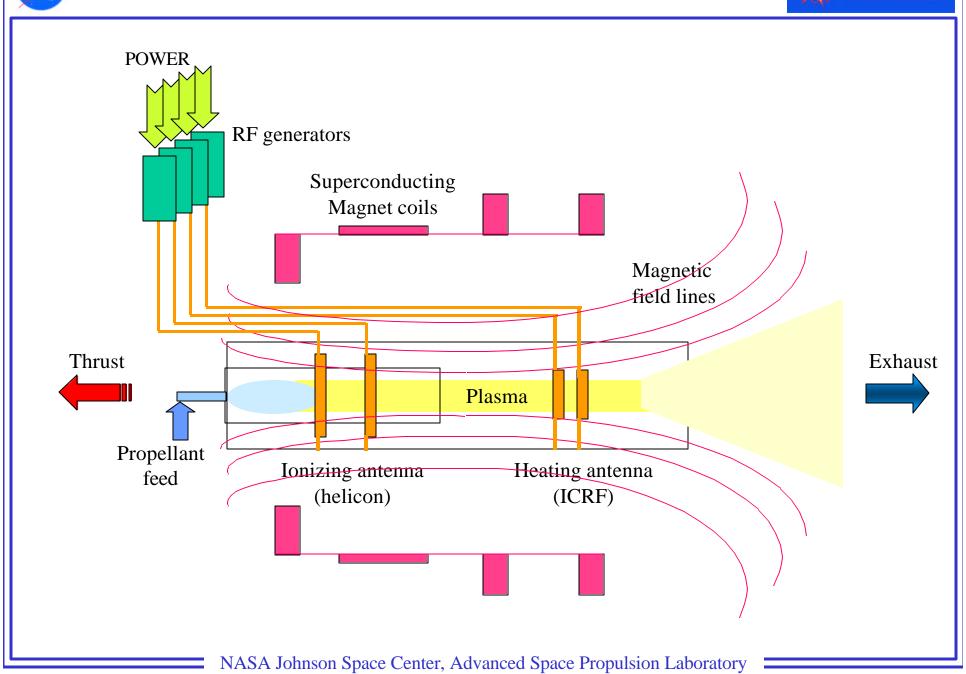
1 Advanced Space Propulsion Laboratory. Johnson Space Center, Houston TX. 2 Dept. of Physics, The University of Texas at Austin, Austin TX. 3 The Oak Ridge National Laboratory, Oak Ridge TN. 4 Massachusetts Institute of Technology, Cambridge MA. 5 Propulsion Research Center, Marshall Space Center, Huntsville AL. 6 University of Houston, Houston TX. 7 Los Alamos National Laboratory, Los Alamos NM. 8 Australian National University, Canberra, Australia.



Simplified Diagram of VASIMR Thruster



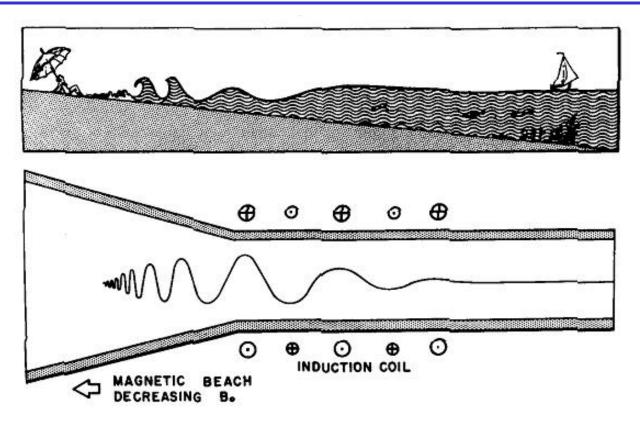
ASPL

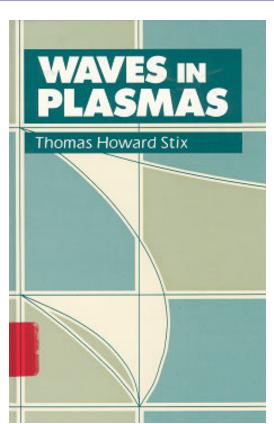




Magnetic beach concept introduced by H. Stix







- •Particles gain energy from the incident wave, like surfers on the ocean
- •Mechanism is proven very efficient in fusion plasmas



Physics of VASIMR



ASPL

Plasma Source

RF waves establish a "helicon" discharge, which ionizes neutral gas to produce a dense plasma with an electron temperature of a few eV

Magnetic Nozzle

When particles see an expanding magnetic field, they are accelerated axially at the expense of their rotational motion.

Ions get extra

(Ion Cyclotron Range of Frequencies)

ICRF heating

Injected electromagnetic waves accelerate the ions by resonating at magnetic beach with their fundamental cyclotron frequency (and associated harmonics.) Ions get extra kick from ambipolar electric field

Both ions and electrons leave at the same rate

NASA Johnson Space Center, Advanced Space Propulsion Laboratory



Important Advantages



- No electrodes or other materials in direct contact with the plasma.
- Therefore, potential for very high power density, high reliability, long life.
- Multiple propellants: Helium, Hydrogen, Deuterium, Nitrogen, Argon, Xenon, others...



NEAR TERM BENEFITS



- DRAG COMPENSATION FOR THE ISS (Hydrogen is a waste gas on the Station)
- SYSTEM BECOMES PROPULSION TECHNOLOGY TEST BED ON THE ISS WITH STRONG COMMERCIAL POTENTIAL
- PLASMA CONTACTOR FOR EVA SAFETY
 - DOD APPLICATIONS



VASIMR Demonstration on ISS



RF Amplifier Set (1 of 4)**Radiator Thruster** Core **Propellant Tanks Approximate Dimensions** Rechargeable of Thruster Core: **Batteries** Diameter < 0.5 mLength < 1 m

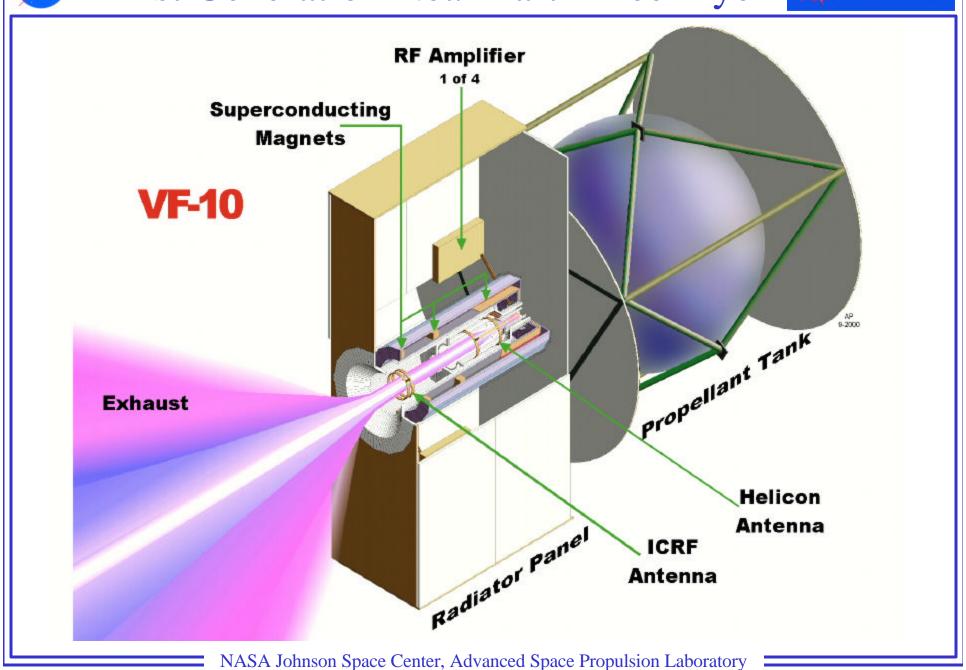
Stepwise Approach:

- Design, build, and test experimental VASIMR thruster in ground test facility and demonstrate performance.
- Test Operation as attached experiment on ISS.
- Use short ~ 10 min firings using stored power (~ 25 kW) from batteries.
- Minimum interfaces with Station



First Generation Near Earth Free Flyer





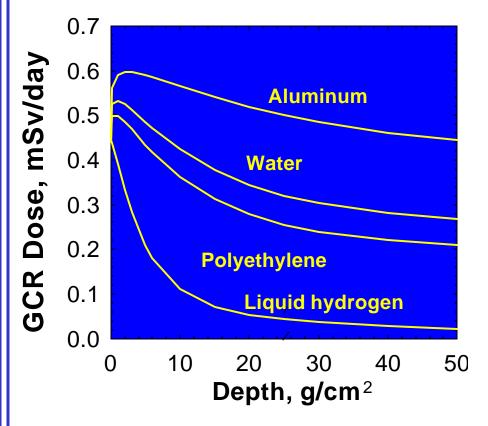


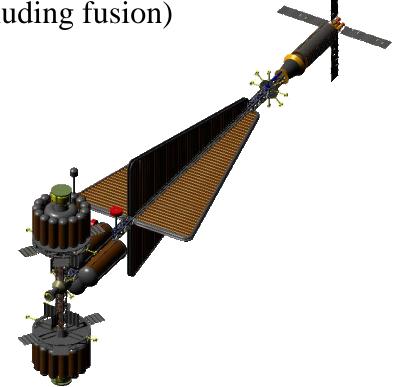
LONG TERM BENEFITS



• Opens entire Solar System to very fast space transportation for humans and robots





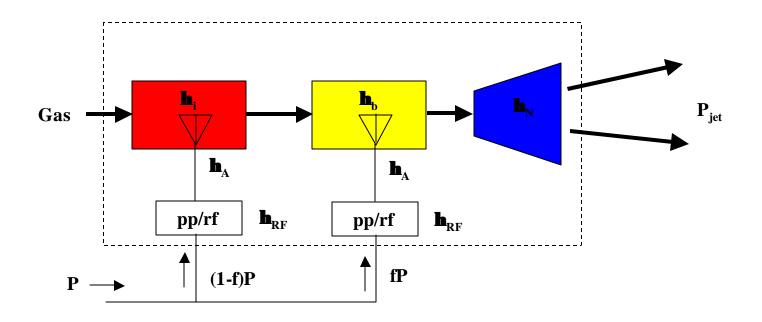


• Hydrogen propellant is plentiful, inexpensive, and best known radiation shield



Engine Efficiency





Definitions:

P DC input power to the PPU in Watts.

 η_{RF} RF power processor unit efficiency

 η_i Helicon efficiency

η_N Nozzle efficiency

Fraction of input power to the RF booster stage

 η_A Transmission line and antenna efficiency

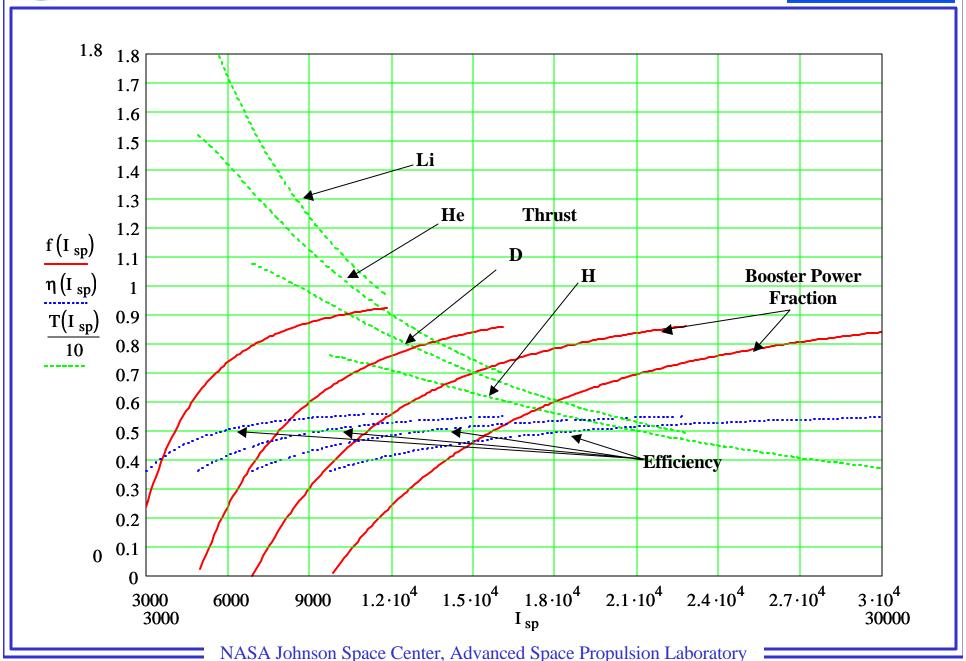
 η_b RF booster efficiency



Engine Performance Estimates with Near-term technology



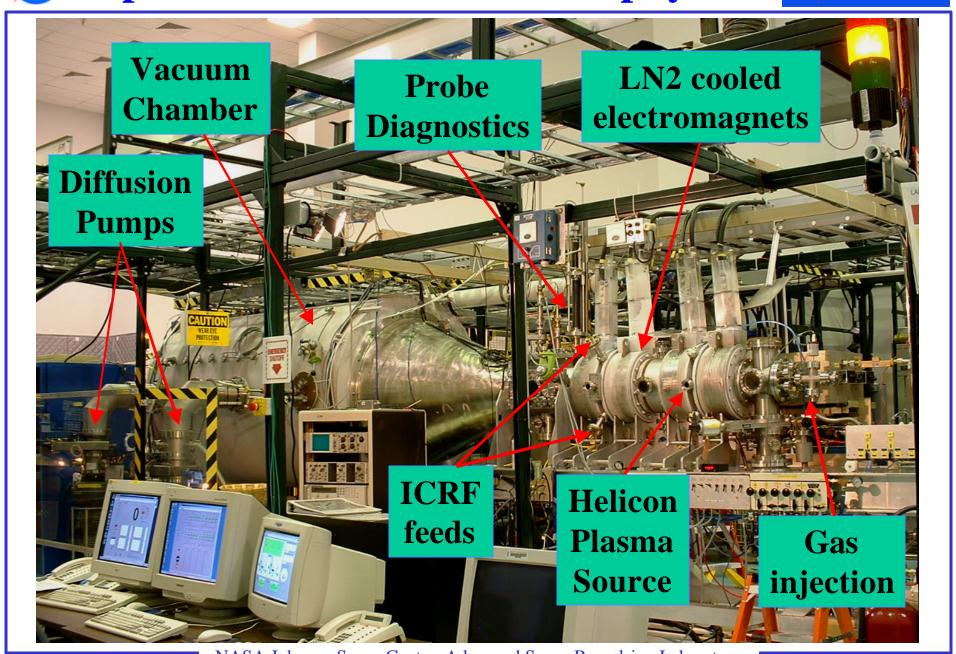
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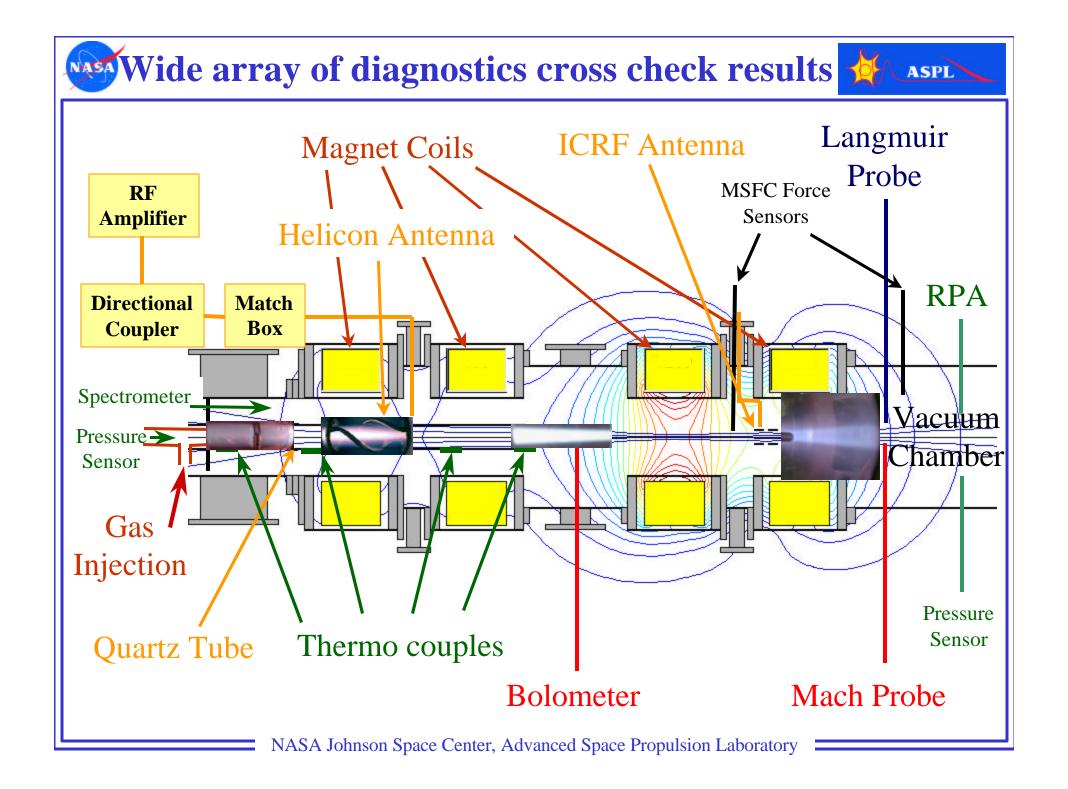




Experiment demonstrates the physics





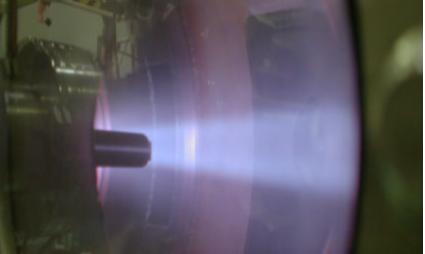




VASIMR Hydrogen Magnetoplasma







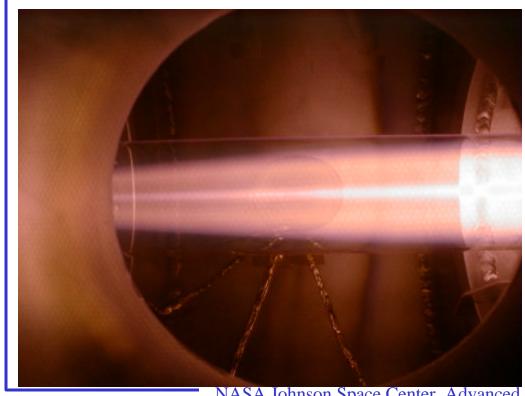


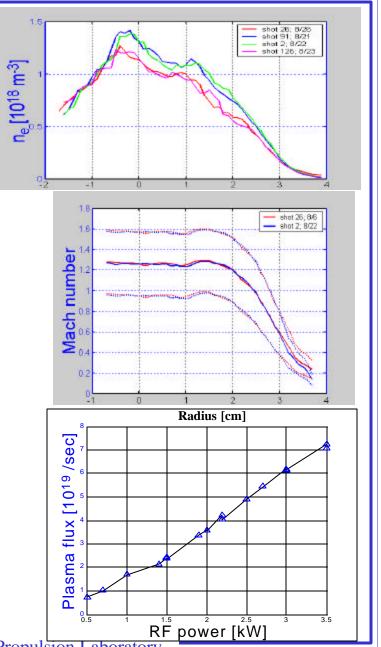


Achieved steady-state operation with light gases



- ➤ High density, stable plasma discharges with many gases are now routine
- **▶**Plasma flows very fast
- ➤ Plasma output linear with input power





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Specific Impulse of Plasma Source



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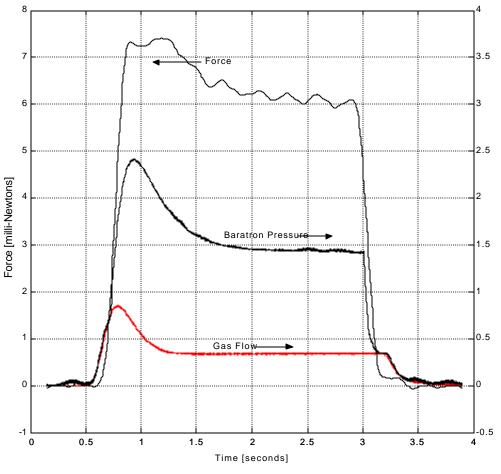
From force measurements, the plasma source ALONE already produces a very high specific impulse. This value is likely to be higher. Measurement is biased low due to pumping limitations.

$$Isp = \frac{T}{\dot{m}g}$$

$$Isp = \frac{T}{\dot{m}g}$$

$$Isp = \frac{.006 \text{ N}}{\left(3 \times 10^{-7} \text{ kg/sec}\right) \left(9.8 \text{ m/s}^2\right)} \begin{bmatrix} \text{Total Markov Ma$$

$$Isp = 2,000 \sec$$

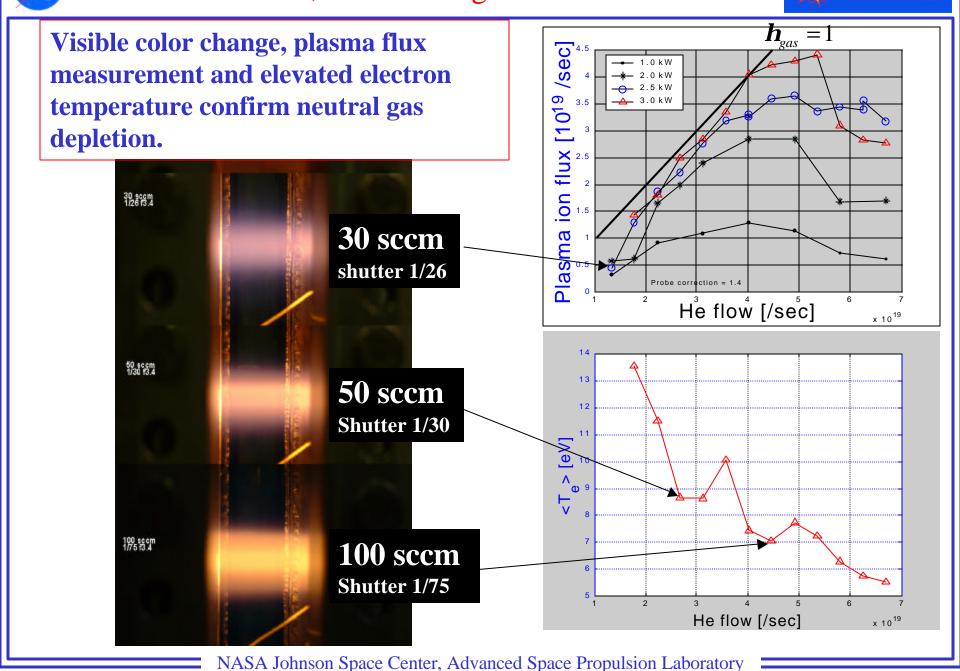






Helium, near 100% gas utilization

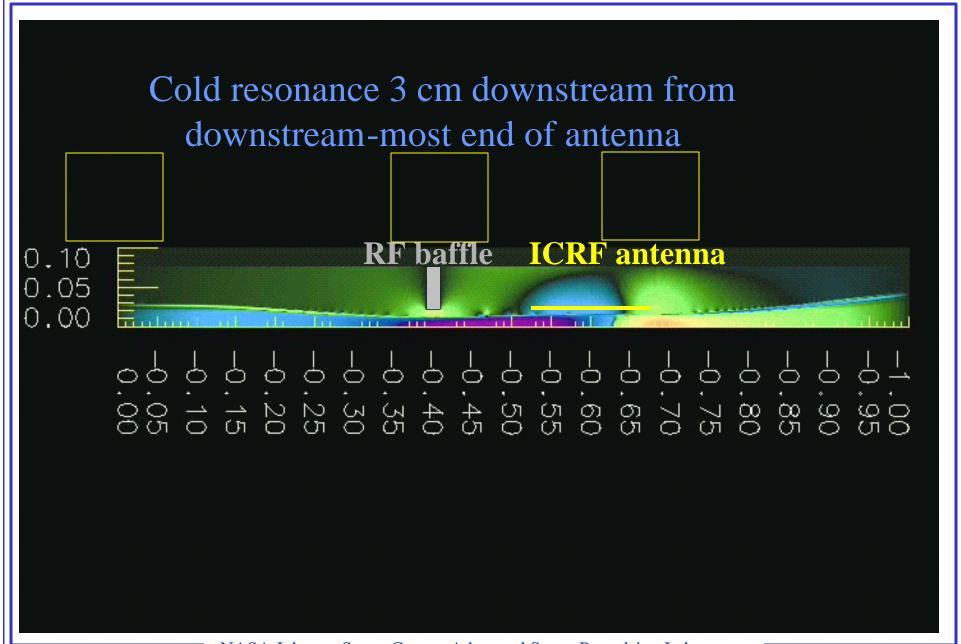






E^+ time evolution

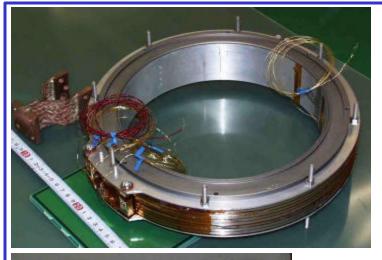






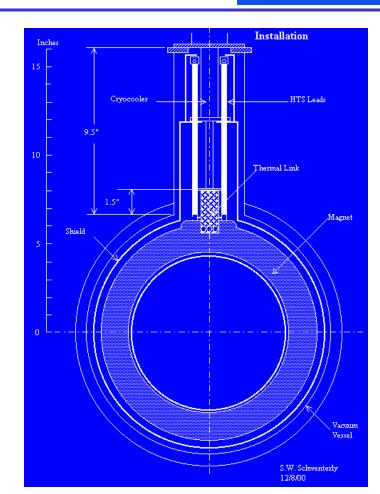
Developed new superconducting magnet technology

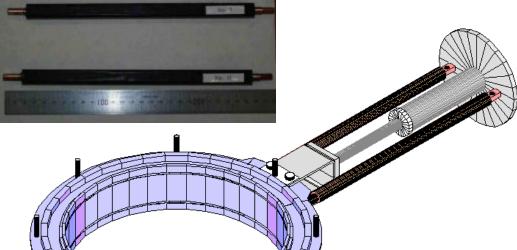




Material:

BSCCO 2223 at 40°K





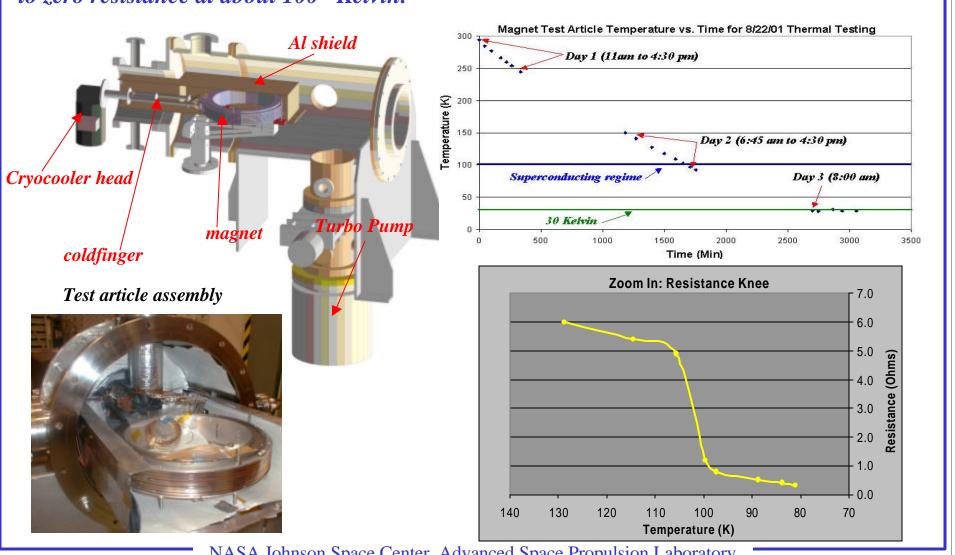
Flight like magnet designed by NASA/ORNL and DuPont Superconductivity Inc. Delivered **April, 2001.**



Advanced Superconducting Magnet Testing



August 9th 2002: ASPL demonstrated high temperature superconductivity by cooling the BSCCO 2223 (flight-like) magnet below critical temperature and observing the transition to zero resistance at about 100 ⁰ Kelvin.



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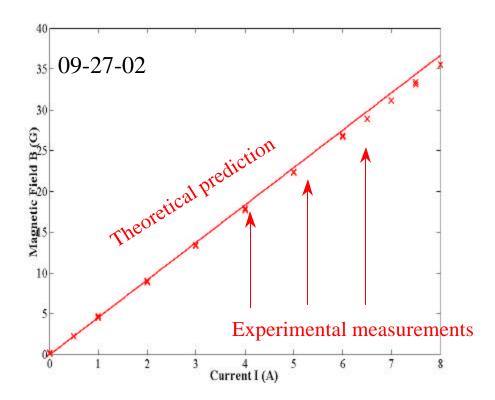


Measured field tracks prediction very well



ASPI

- Magnet was operated on September 24 and 27, 2002 with up to 8 amps.
- Expected magnetic field strength was measured
- Temperature was stable during testing
- Maximum current tested so far is 79 amps, goal is 105 amps
- Cryocooler availability is pacing item



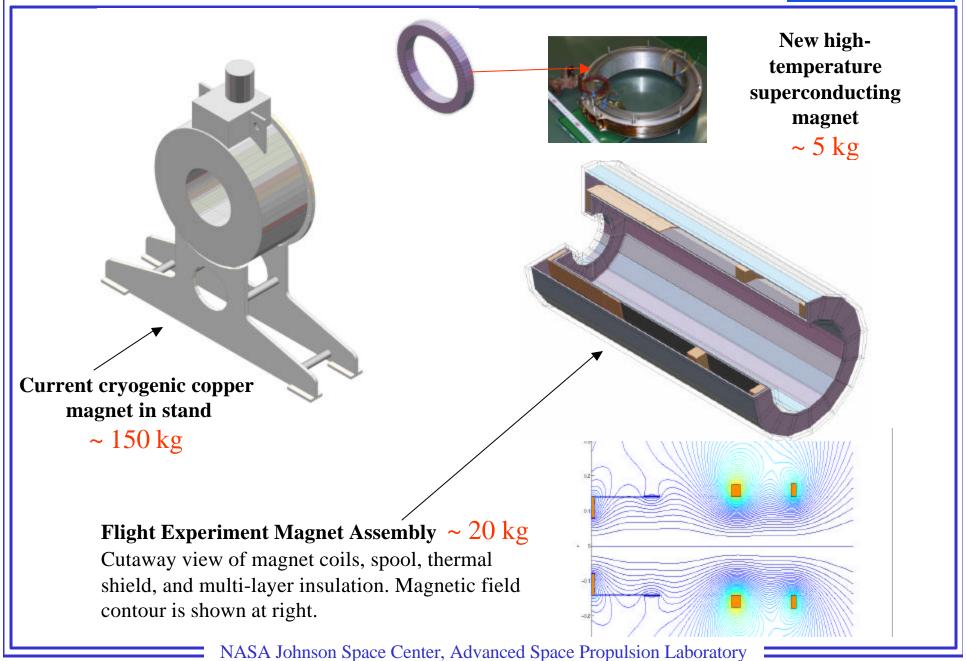




New magnet brings dramatic mass reduction



ASPL

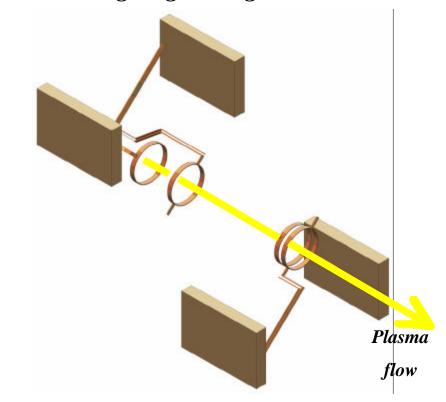


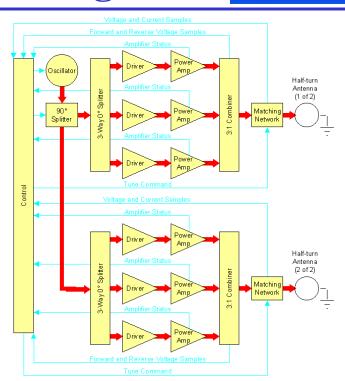


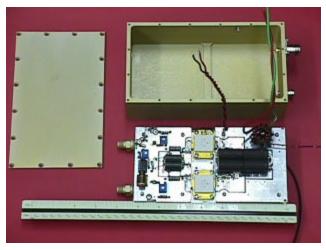
Solid state RF system design



- Design draws from ORNL expertise in RF heating of fusion plasmas.
- System architecture is robust and failure tolerant.
- Prototype hardware has been built and is undergoing testing.







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Collaboration with DOE Labs



ASPL

The DOE laboratories have gathered a lot of experience through decades of research in magnetic confinement fusion.

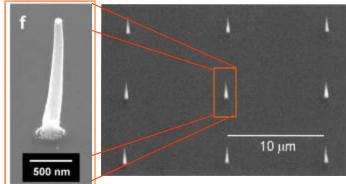






Multi Megawatt antennas for plasma heating are already operational

Carbon nanotube technology for cold, field emission cathodes will increase efficiency





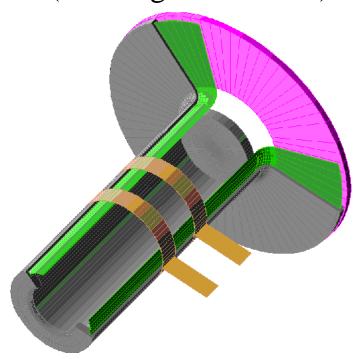
We are developing advanced thermal strategies



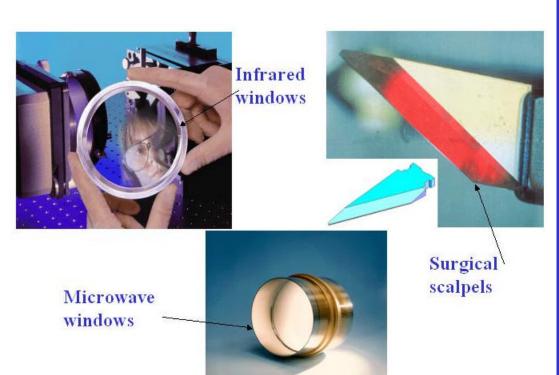
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• Remove heat nearest its source at high temperature

Integrated heat pipes and Cryocooler technology (working with GSFC) Advanced materials (working with industry)



Helicon heat pipe



CVD Diamond Technology from ManSat Inc.

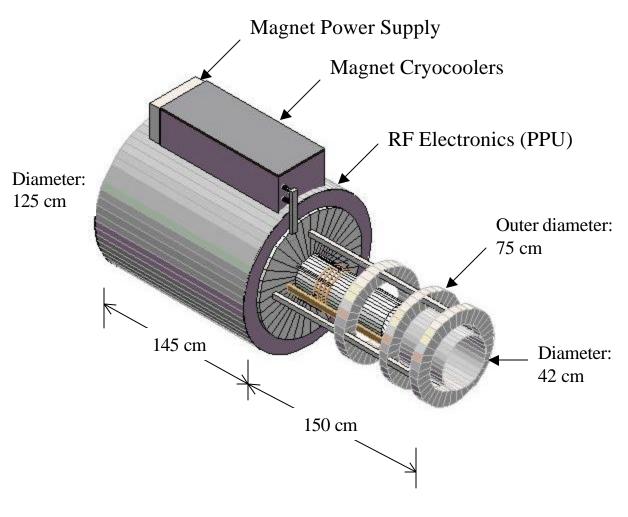


Completed Point Design for One MWatt Engine



ASPL

Engine, PPU and Cryocoolers





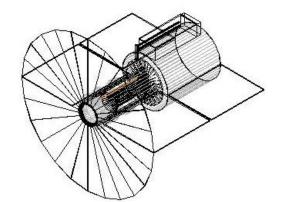
Conservative mass estimate very attractive



Engine	Mass	Dimensions		
	(kg)	(cm)		
Propellant controller	5	30 x 20 x 10		
Engine tube	5.4	55 x 42 dia .4 thick		
Helicon tube	8.5	85 x 30 dia .3 thick		
Helicon antenna	0.27	1.3 x 30 dia .2 thick		
ICRF antenna	1.2	4.5 x 20 dia .2 thick		
Helicon transmission lines	0.2	2 x.5 x 75		
ICRF transmission lines	0.4	2 x .5 x 150		
Magnet power supply	10	45 x 30 x 10		
Magnet cryocoolers (3)	55	115 x 45 x 30		
Magnet loop heat pipe	3	300 x 3		
Cryocooler radiator	2.2	22 x 115		
Magnet coils (3)	70	55 ID 60 OD 5 thick		
Magnet support and insulation	10.5			
Instrumentation	5			
System controller	9	26 x 50 x 9		
Engine radiator	124	3.64 m dia + 2x(1.38 x 1.5)		
Engine support structure	31			
Engine Total	340.67			

Power Processing Unit	Mass	Dimensions	
	(kg)	(cm)	
RF power distribution	150	2x(140 x 28 x 10)	
Helicon oscillator (4)	4.52	4x(15 x 8 x 5)	
Helicon driver (4)	5.2	4x(20 x 20 x 8)	
Helicon power amplifier (4)	144.8	4x(106 x 23 dia)	
Helicon tuned line matcher (4)	21.6	4x(143 x 25 dia)	
Shorted stub matcher (4)	21.6	4x(143 x 25 dia)	
ICRF oscillator (4)	4.52	4x(15 x 8 x 5)	
ICRF driver (4)	60	4x(21 x 21 x 8)	
ICRF power amplifier (4)	160	4x(108 x 23 x 8)	
ICRF matching network (4)	100	4x(54 x 12 x 4)	
ICRF Antenna tuner (4)	120	4x(54 x 12 x 5)	
PPU radiator	45	2x(1.2 x 1.45) + 2.8 m2	
PPU structure and fittings	84		
PPU Total	921.24		

Radiator Panel Mass: One-sided: $4.9~{\rm kg/m^2}~{\rm Two}$ -sided: $8.9~{\rm kg/m^2}$ Antenna sizing based on $5~{\rm MW/m^2}$



Mass Estimate for 1 MW Point Design

Engine: $341 \text{ kg} \quad \alpha = 0.3 \text{ kg/kWe}$

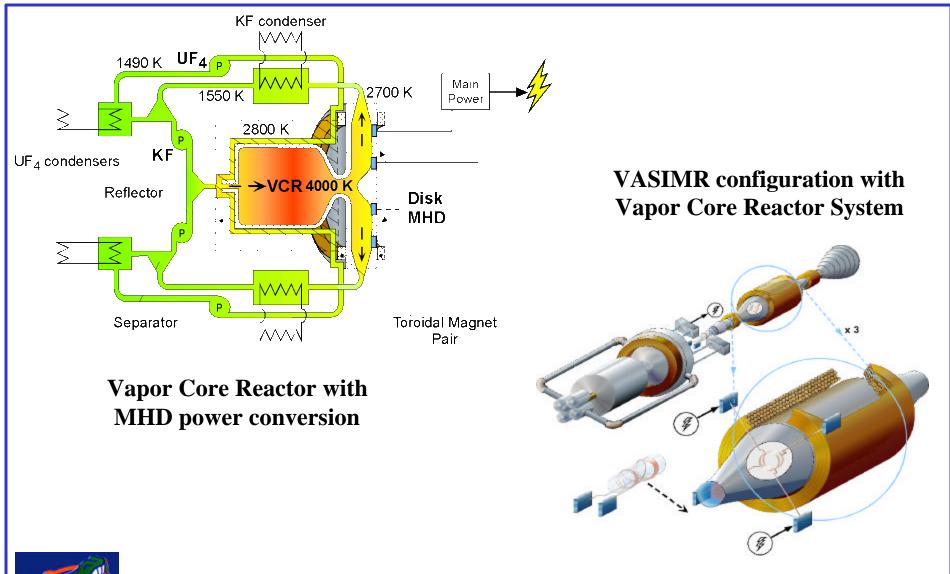
PPU: $921 \text{ kg} \quad \alpha = 0.9 \text{ kg/kWe}$

Total System: $1262 \text{ kg} \quad \alpha = 1.26 \text{ kg/kWe}$



Development of Multi MW Nuclear Power Systems





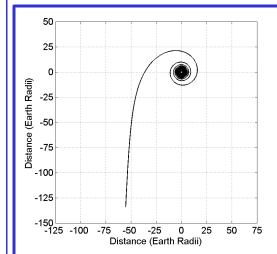


Prof. Samim Anghaie, Director, Innovative Nuclear Space Power and Propulsion Institute, INSPI; University of Florida, Gainesville.



Fast (115day) Mars Mission Architecture





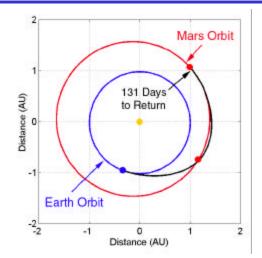
High thrust Earth spiral (30days)

Departing LEO May 6, 2018

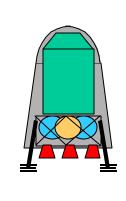
188 mT IMLEO

12 MW power plant

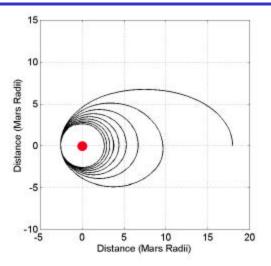
a = 4 kg/kW



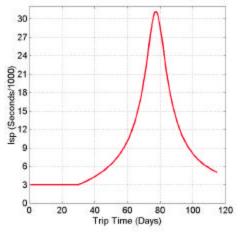
Heliocentric Trajectory(85days)



Crew Lander (60.8 mT Payload) 31.0 mT Habitat 13.5 mT Aeroshell 16.3 mT Descent System



Robotic Mars orbit insertion



Isp profile for piloted segment



Higher Power dramatically reduces trip time



200MW Earth to Mars Missions $\alpha = 0.5$; Maximal $I_{sp} = 30,000$ Payload Mass 22 mT

Total Initial			Heliocentric trajectory fuel (mT) time (days)		Final relative	Total trip
Mass (mT)					velocity (km/s)	time (days)
600	180	7	298	34	0	41
350	117	5	111	42	0	47
250	88	4	40	49	0	53
600	152	8	324	31	6.8	39



VASIMR enables contingency abort capability

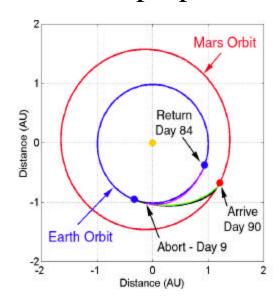


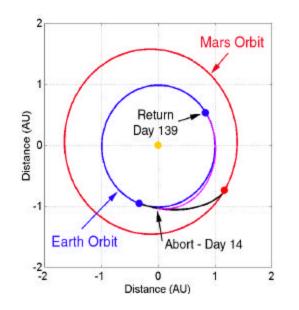
ASPL

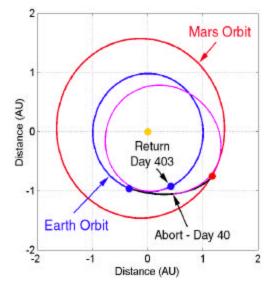
With variable I_{sp} , operational flexibility is increased in the event of loss of propellant or other system failures

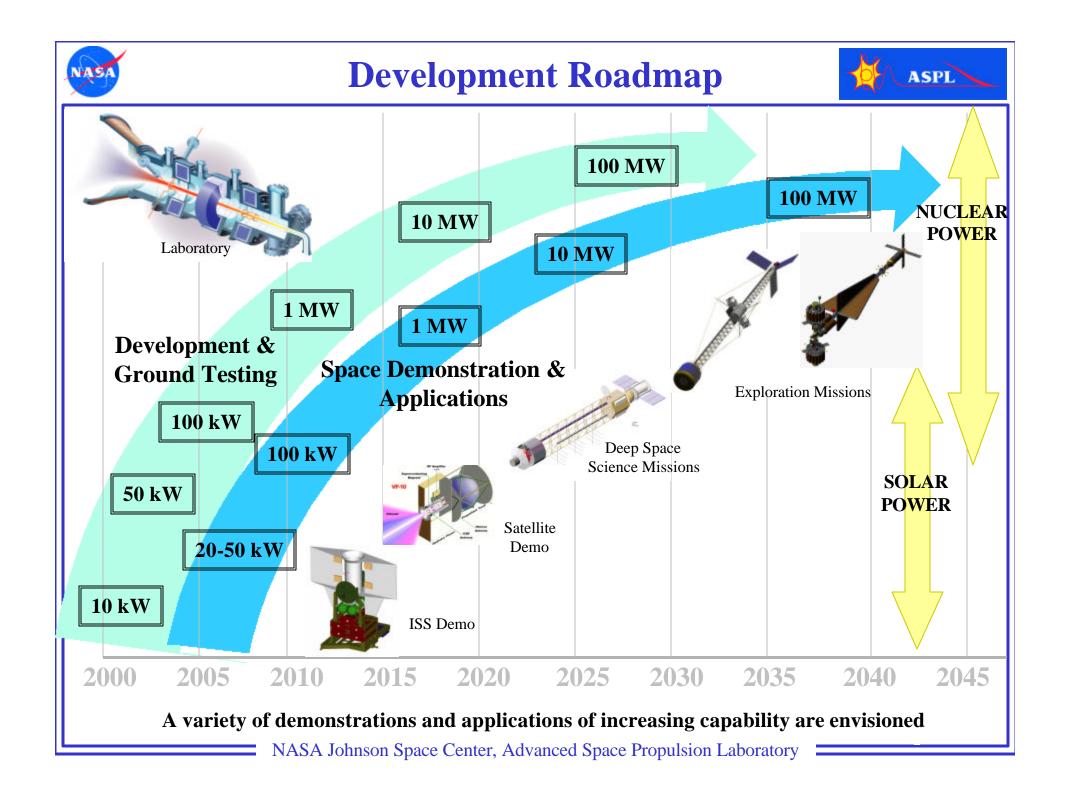
Aborts due to loss of propellant

Aborts deep into the mission due to non propulsion system failures











Main DOE and University Collaborators



- Oak Ridge National Laboratory, Fusion Energy Division: Dr. Wally Baity, RF systems
 - Dr. Mark Carter, RF systems, plasma theory, magnetic system design
 - Dr. Rick Goulding, experimental plasma generation and heating
 - Dr. William Schwenterly, superconducting magnet design
- Los Alamos National Laboratory: Dr. Patrick Colestock, wave physics and simulation
 - Dr. Max Light, helicon physics and wave diagnostics
- University of Alabama, Huntsville: Dept of Physics: Dr. James Miller
 - Mr. Greg Chavers, plume energy and momentum measurements (Also, from NASA, MSFC)
- Australian National University: Plasma Research Laboratory: Dr. Roderick Boswell, Helicon design and wave physics
 - Dr. Christine Charles, magnetized plume physics
 - Mr. Orson Sutherland, helicon physics
- University of Texas, Austin, Fusion Research Center: Dr. Roger Bengtson, experimental plasma physics and diagnostics
 - Dr. Boris Breizman, plasma theory and system scaling
 - Dr. Alexei Arefiev, plasma theory and system scaling
 - Dr. Cesar Ocampo, Trajectory design and optimization
 - Mr. Christopher Rainieri, trajectory design and optimization
- Costa Rica Center for High Technology
 - Dr. Jorge Andrés Díaz, mass spectroscopy and recombination chemistry
- University of Florida at Gainesville, Inovative Space Nuclear Power Institute
 - Dr. Samim Anghaie (Director), Space Nuclear Reactor Design
- Rice University, Dept. of Physics and Astronomy: Dr. Anthony Chan, plasma theory
 - Dr. Carter Kittrell, experimental plasma spectroscopy
 - Dr. Timothy Glover, plasma diagnostics, optical interferometry
- University of Houston, Dept. of Physics:
 - Dr. Edgar Bering, experimental plasma physics and ion diagnostics
- Alfven Laboratory, Swedish Royal Institute of Technology:
 - Dr. Nils Brenning, RF wave physics (experiment)
 - Dr. Einar Tenfors, wave physics (theory)
- MIT, Plasma Science and Fusion Center: Dr. Kim Molvig, Plasma theory and simulation
 - Dr. Oleg Batischev, plasma non-linear theory and simulation
- University of Michigan: Dr. Brian Gilchrist
 - Mr. Christopher Davis, plasma interferometry



VASIMR Workshop 2002



